

Redshift Evolution of the Electron Density and  
Extremely Low C/N Galaxies with supersolar N/O at  $z \sim 4-10$   
Uncovered with JWST/NIRSpec

Isobe+23a, arXiv: 2301.06811

Isobe+23b, arXiv: 2307.00710

*Released yesterday!*

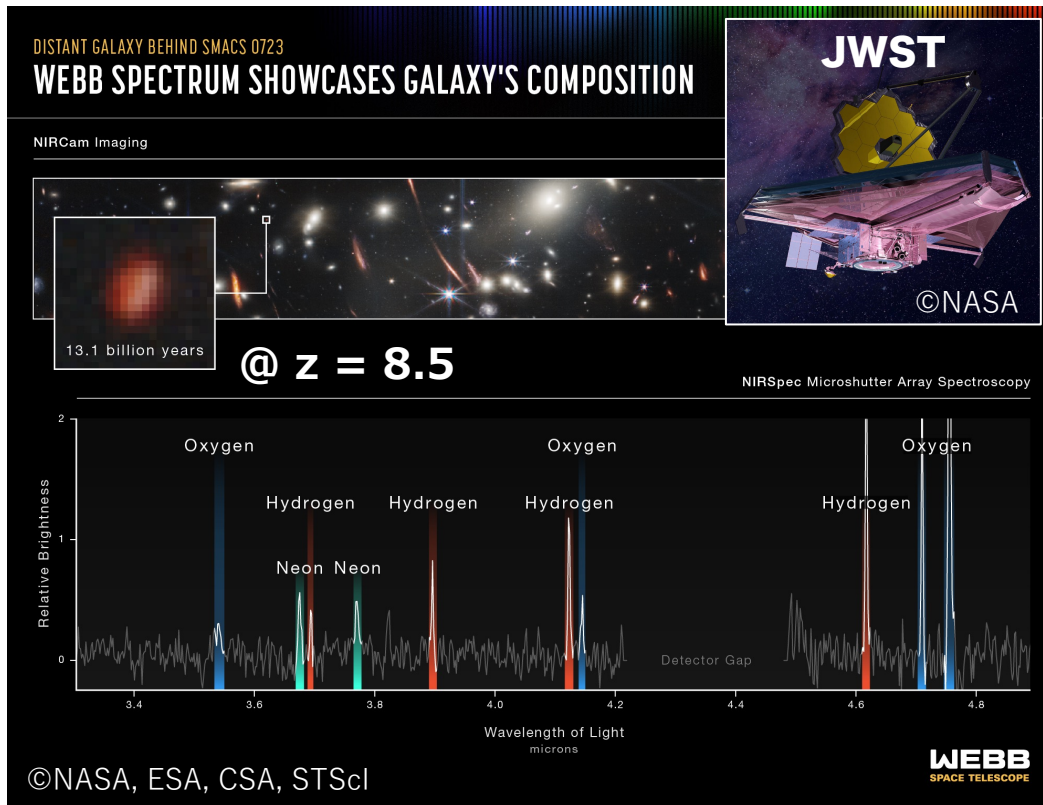
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Collaborator: [M. Ouchi](#), N. Tominaga, K. Watanabe, [K. Nakajima](#), H. Umeda,  
H. Yajima, [Y. Harikane](#), H. Fukushima, Y. Xu, Y. Ono, and Y. Zhang



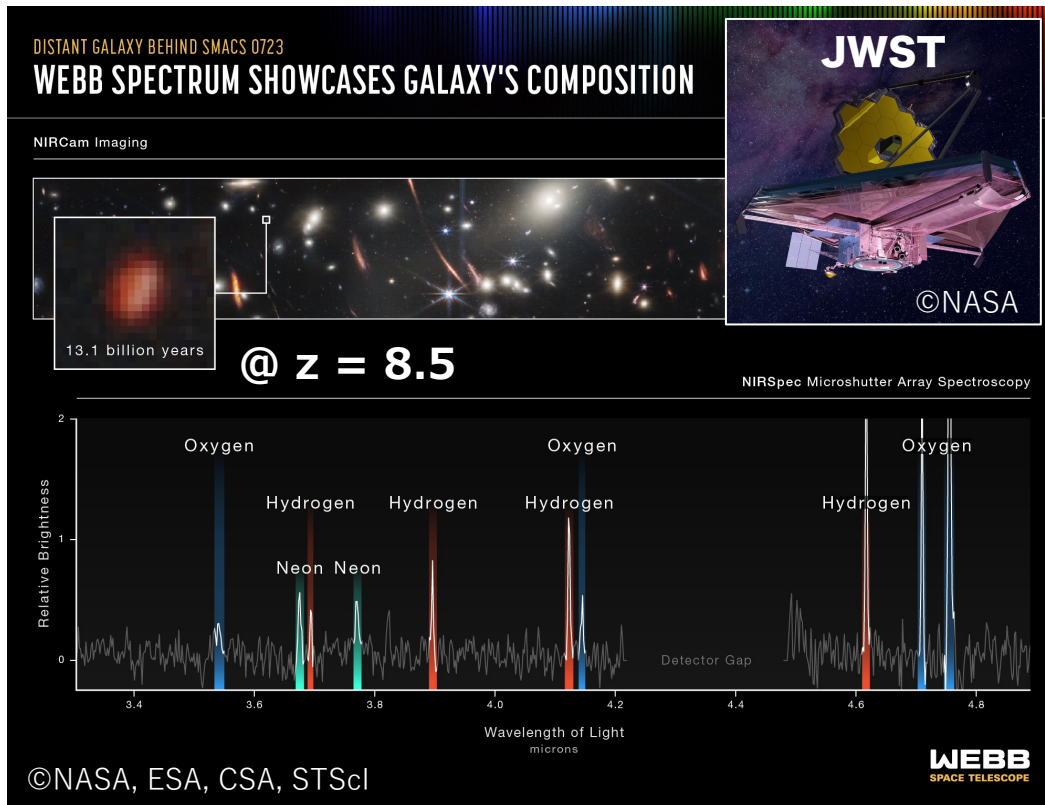
# Introduction

## Amazing JWST/NIRSpec spectra

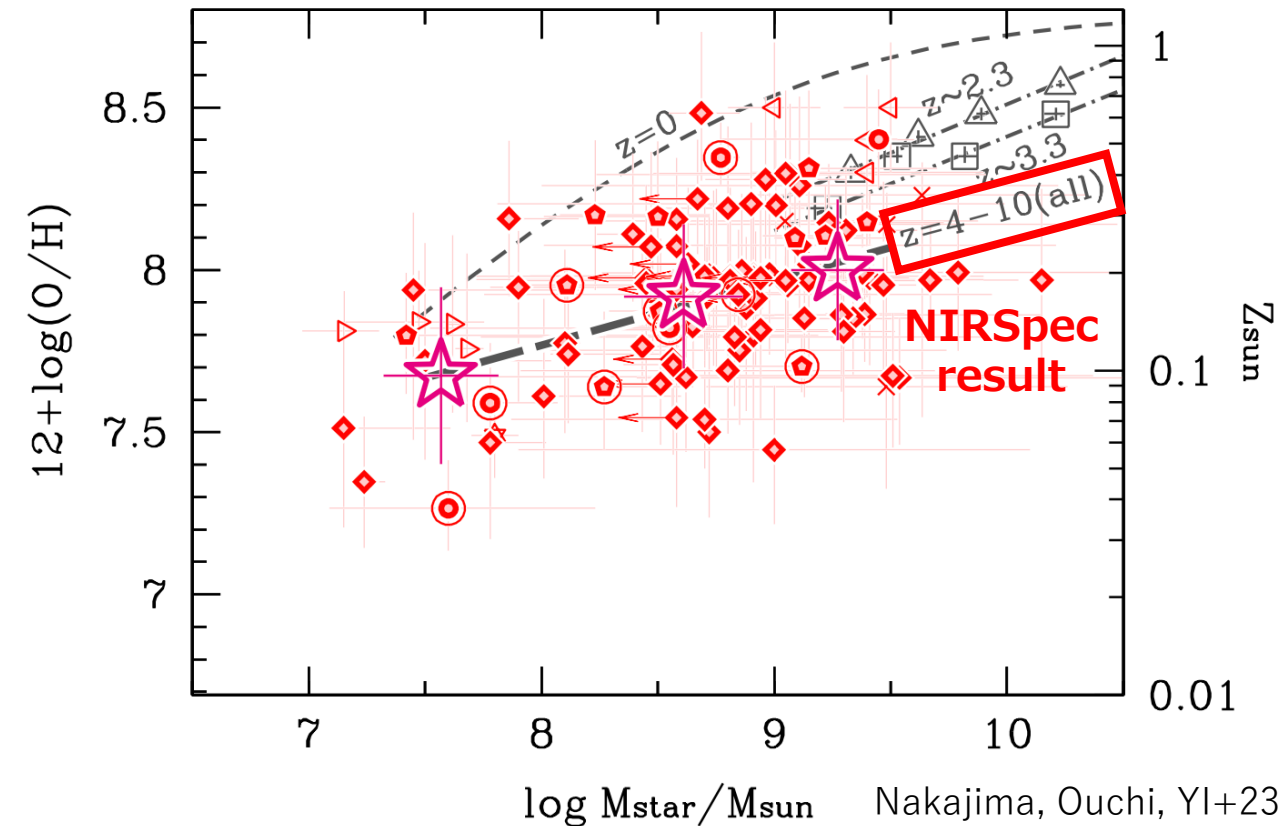


# Introduction

## Amazing JWST/NIRSpec spectra



## E.g., Metallicity measurement at $z \sim 4-10$ (Nakajima+23; Curti+23; ...)



# Data and Sample

Thanks to the NIRSPEC public surveys

ERO (Pontoppidan+22)

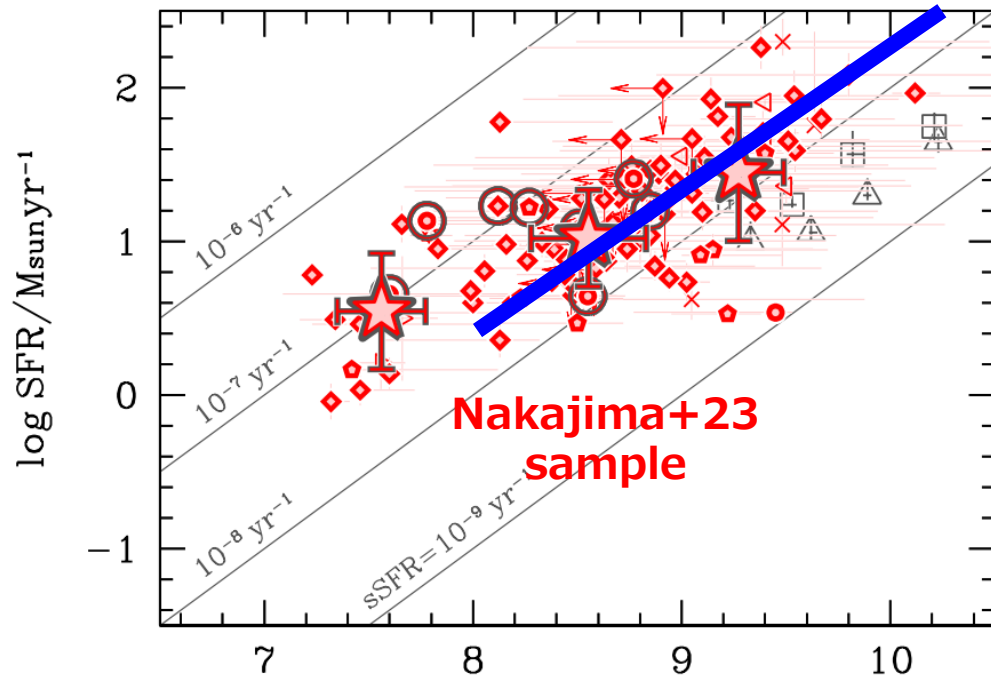
GLASS (Treu+22)

CEERS (Finkelstein+22)

Data reduction by Nakajima+23

→ 111 galaxies at  $z = 4-10$

$z = 5-6$  main sequence  
(Santini+17)



log Mstar / Msun Nakajima, Ouchi, YI+23

# Data and Sample

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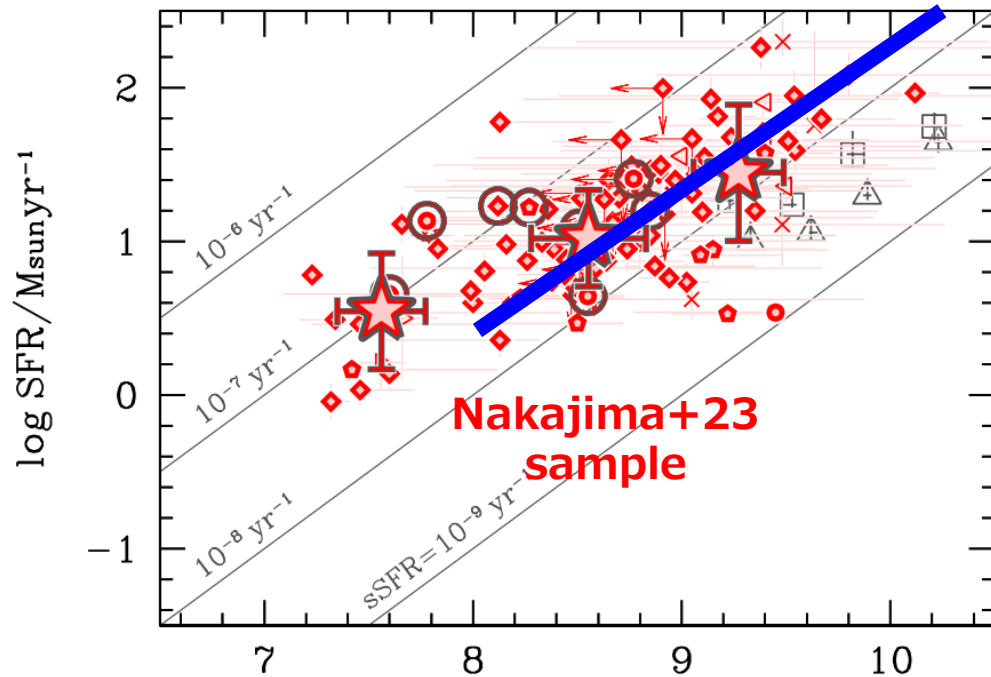
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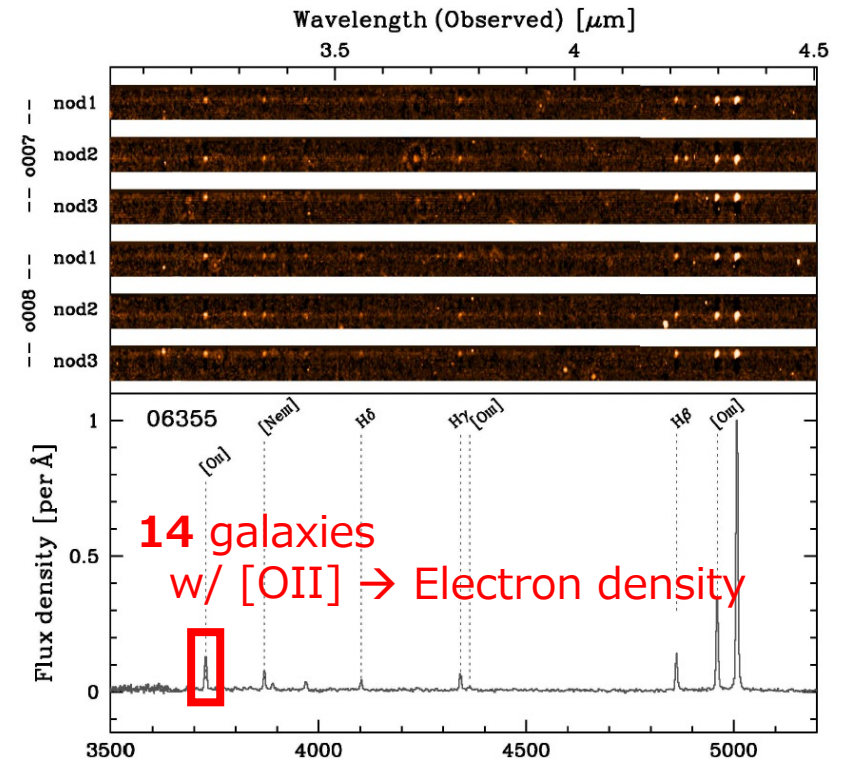
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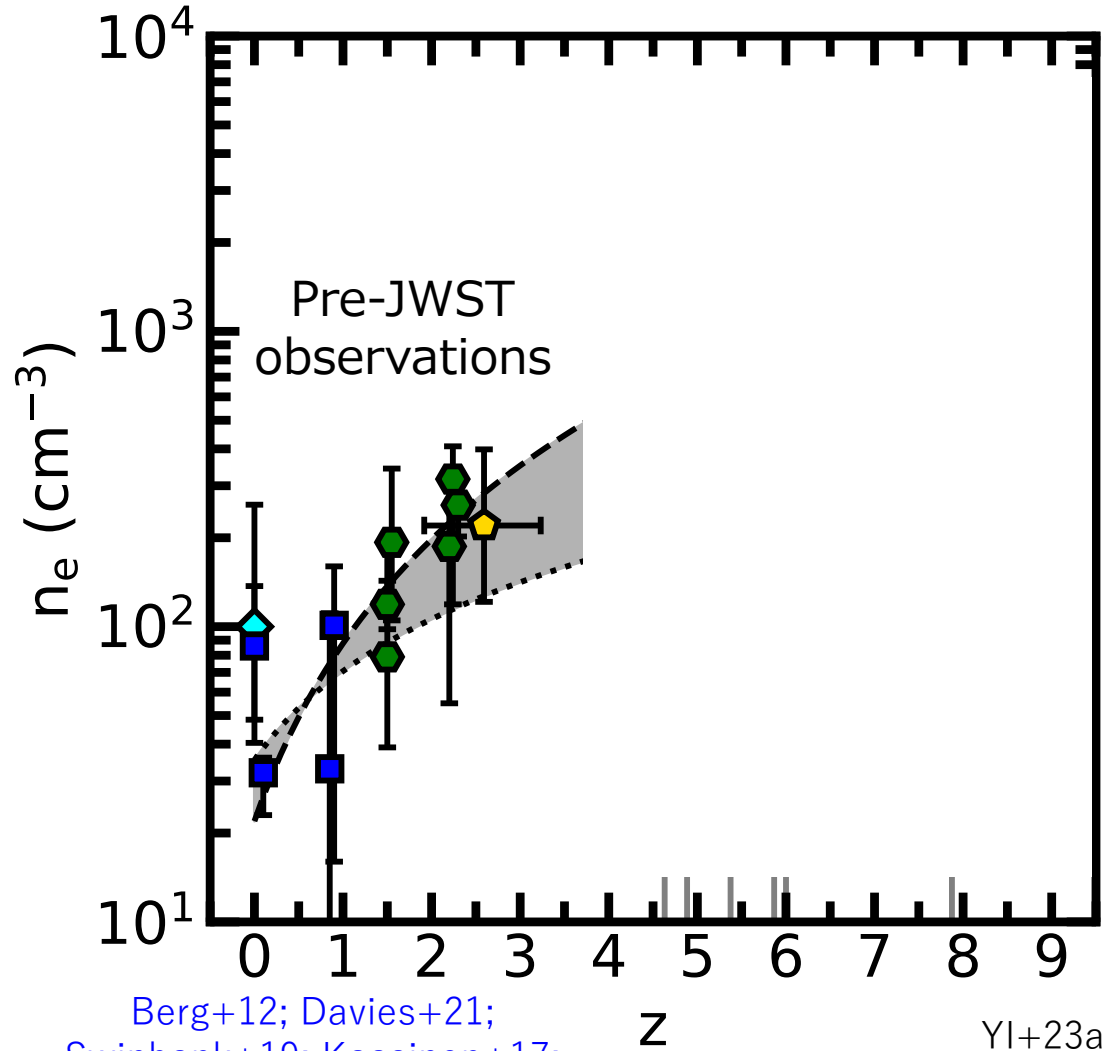
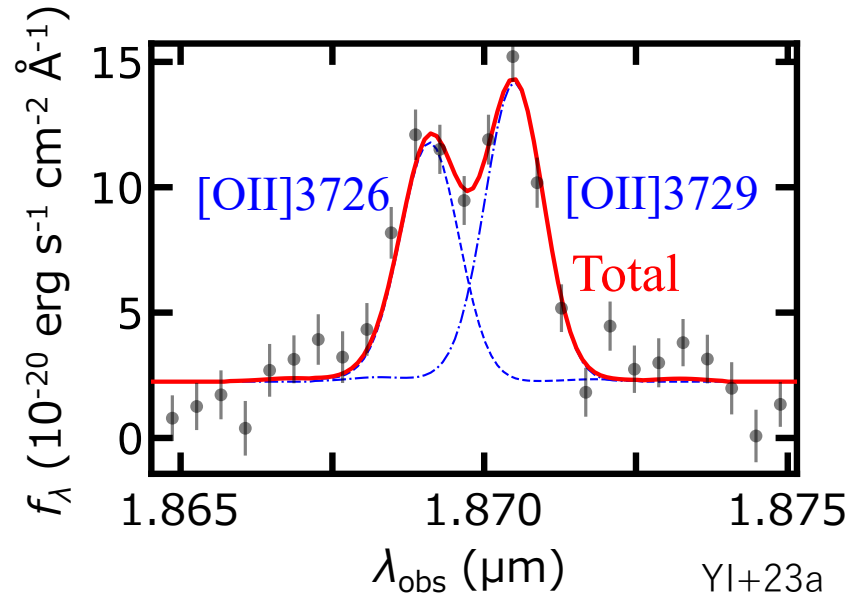
$\log M_{\text{star}}/M_{\text{sun}}$  Nakajima, Ouchi, YI+23



Wavelength (Rest) [ $\text{\AA}$ ] Nakajima, Ouchi, YI+23

# Redshift Evolution of Electron Density

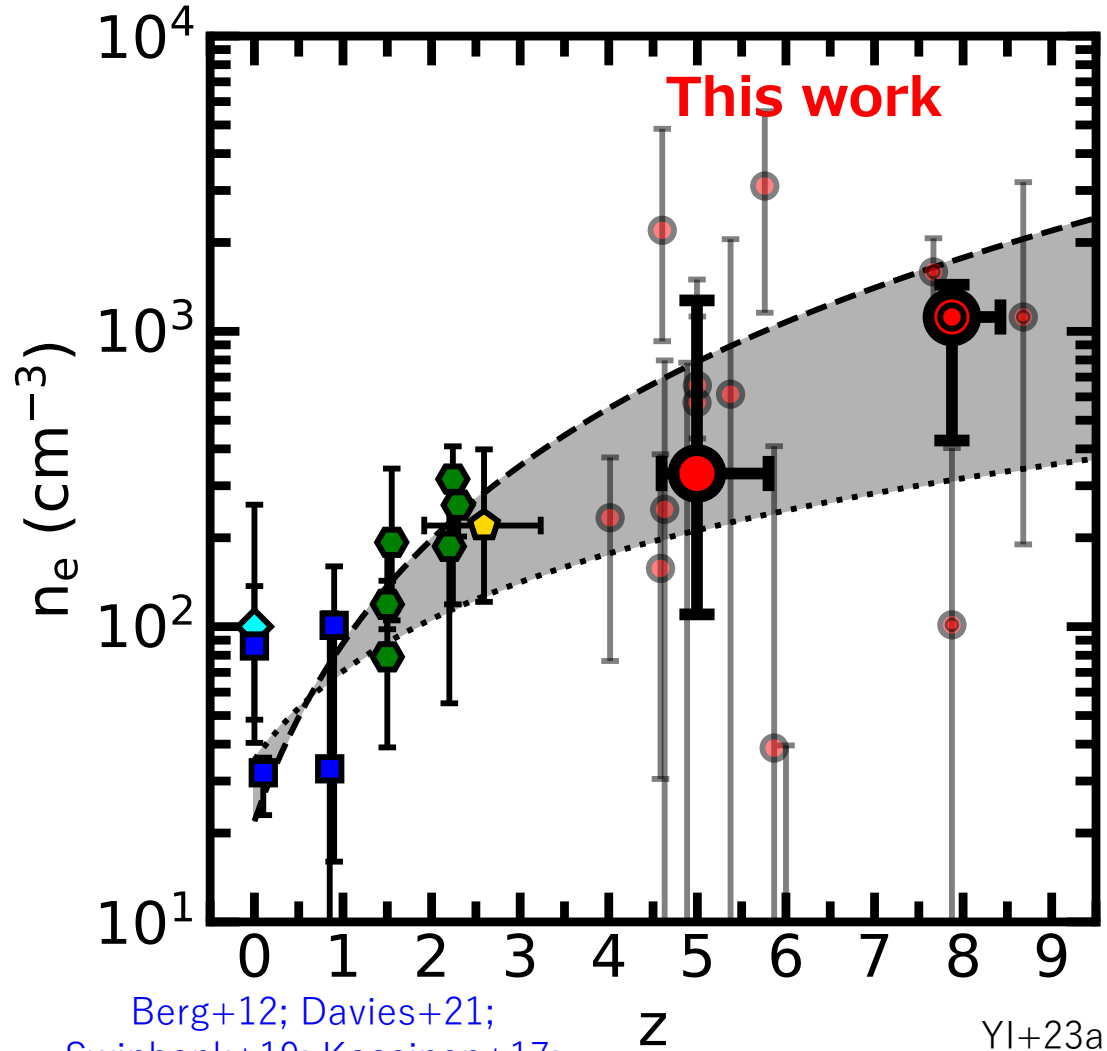
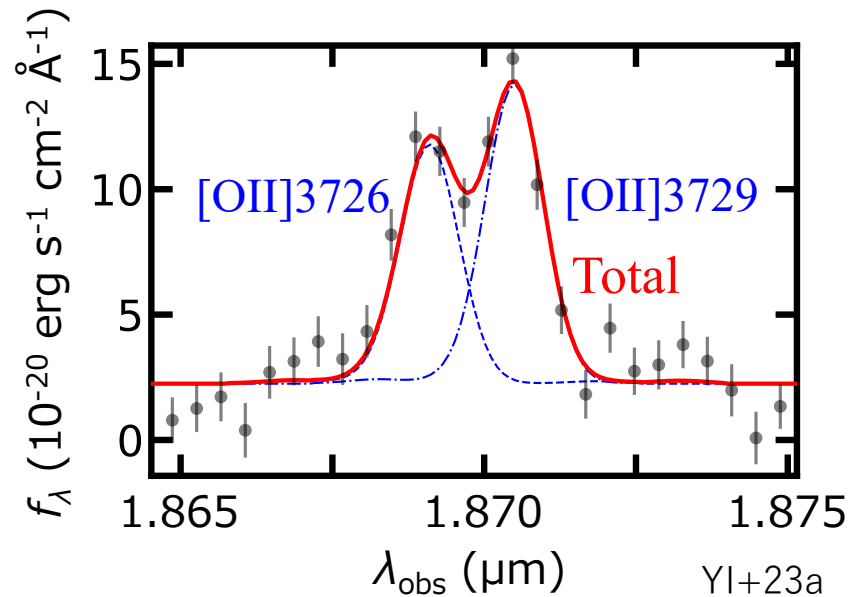
Deriving LSFs for  
careful [OII] separation



Berg+12; Davies+21;  
Swinbank+19; Kaasinen+17;  
Davies+21; Kashino+17;  
Sanders+16; Steidel+14

# Redshift Evolution of Election Density

Deriving LSFs for  
careful [OII] separation

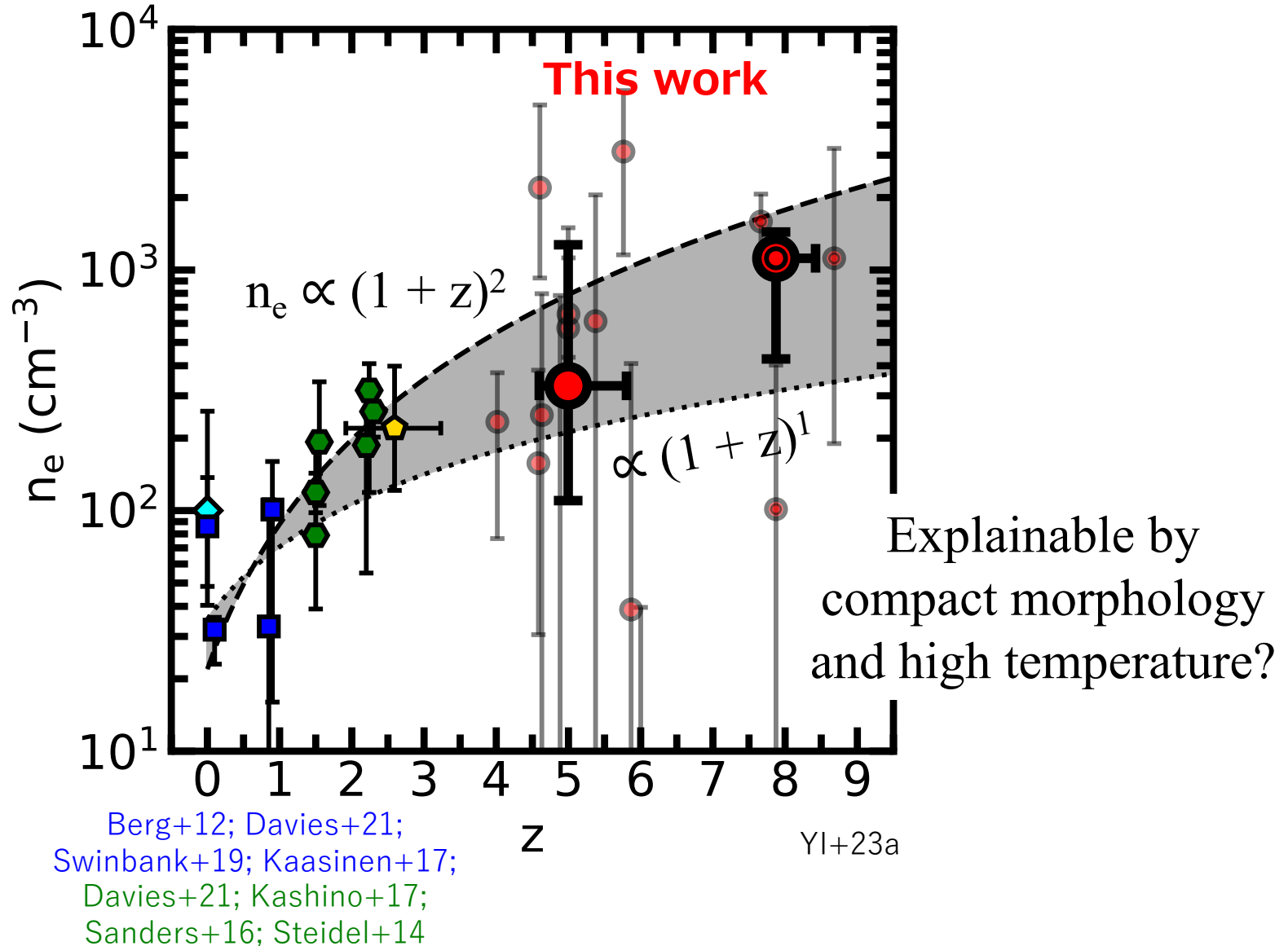
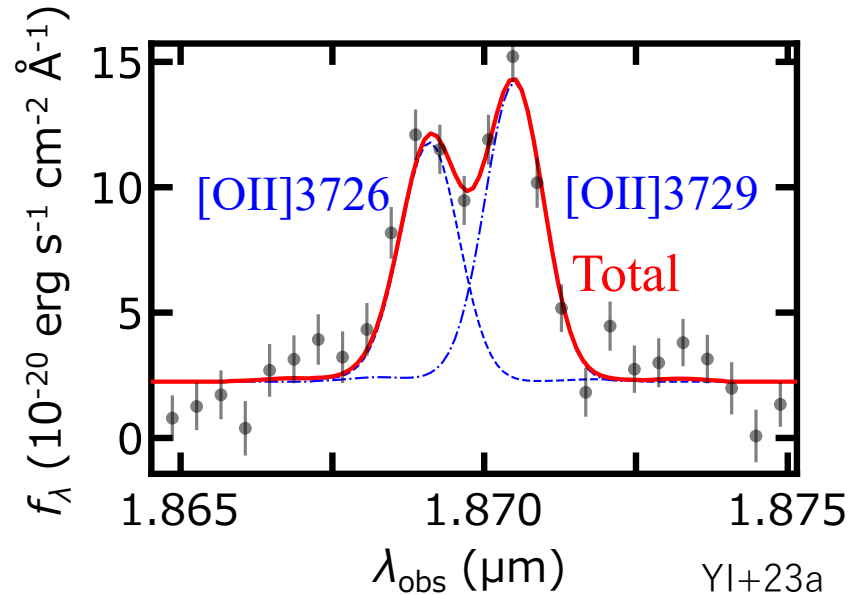


**Increase!**

Berg+12; Davies+21;  
Swinbank+19; Kaasinen+17;  
Davies+21; Kashino+17;  
Sanders+16; Steidel+14

# Redshift Evolution of Election Density

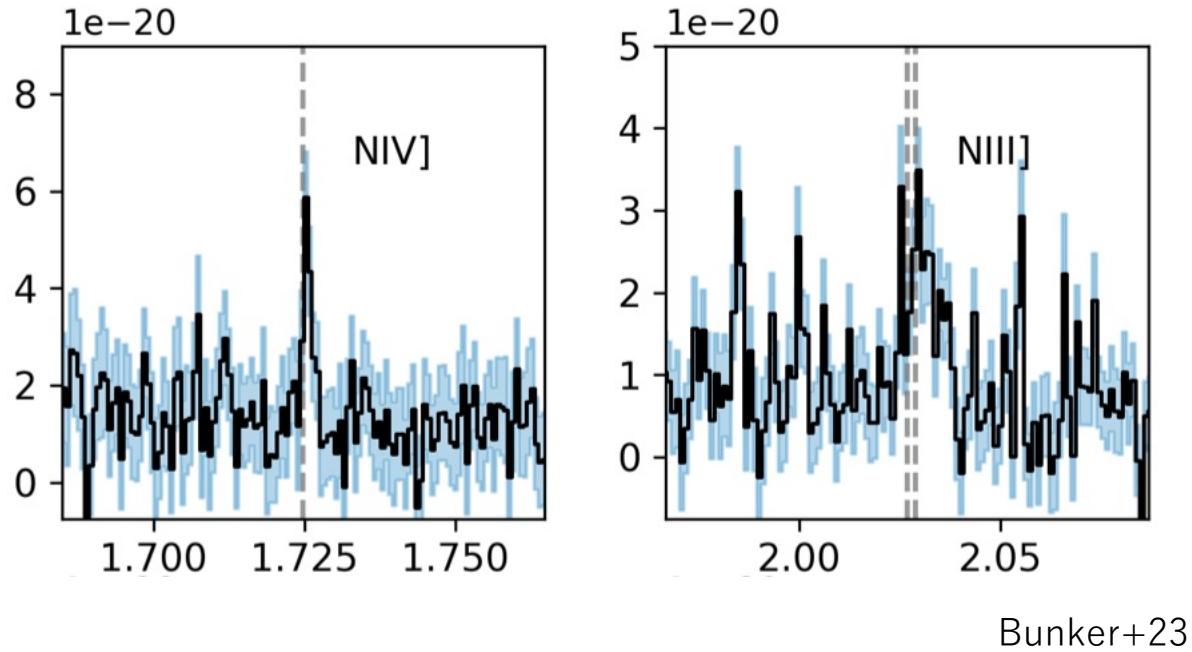
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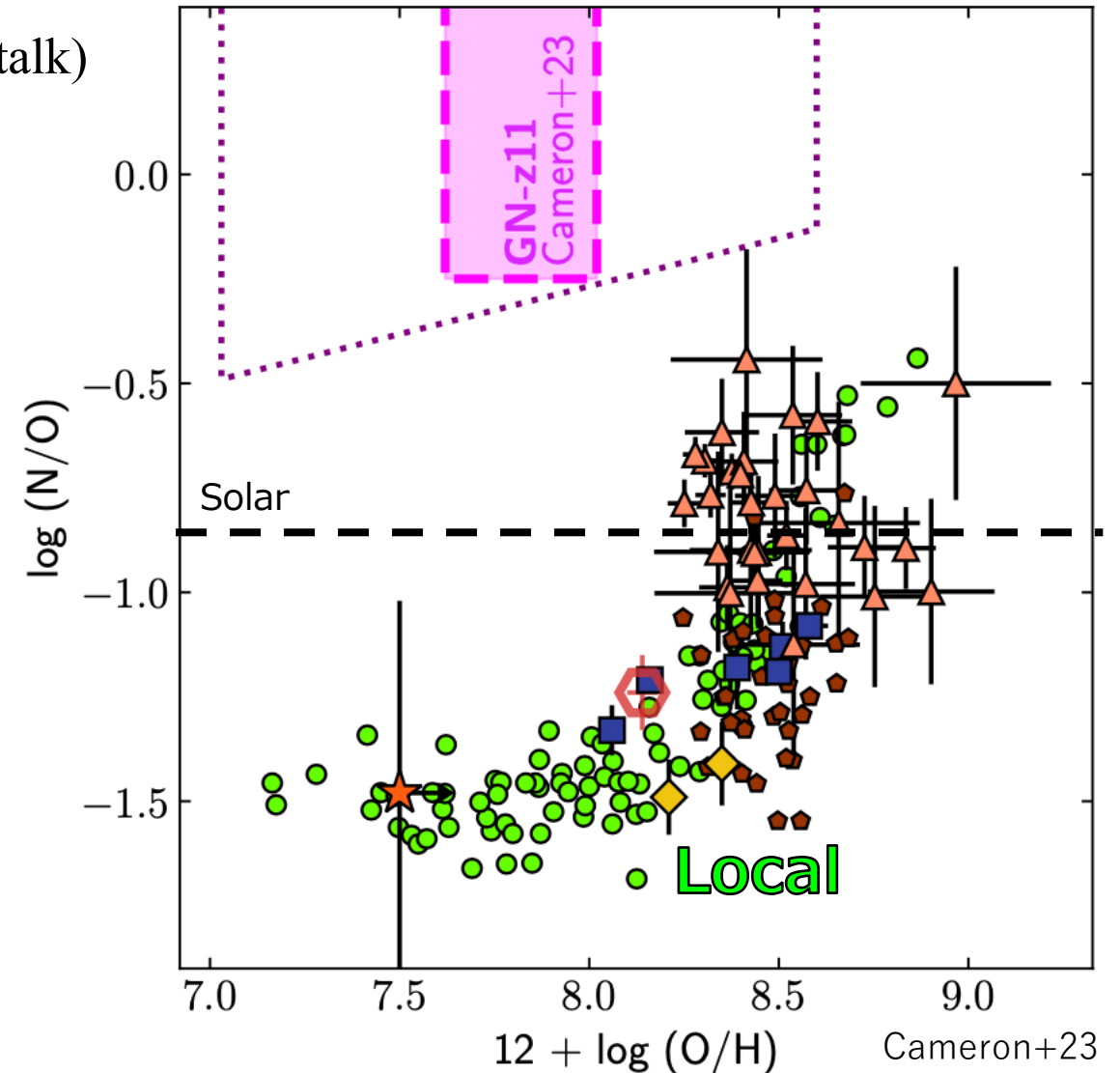


# Supersolar N/O in GN-z11 @ $z = 10.6$

Unexpectedly strong NIII] and NIV] lines  
from GN-z11 (JADES; Bunker+23; see Jan Scholtz's talk)

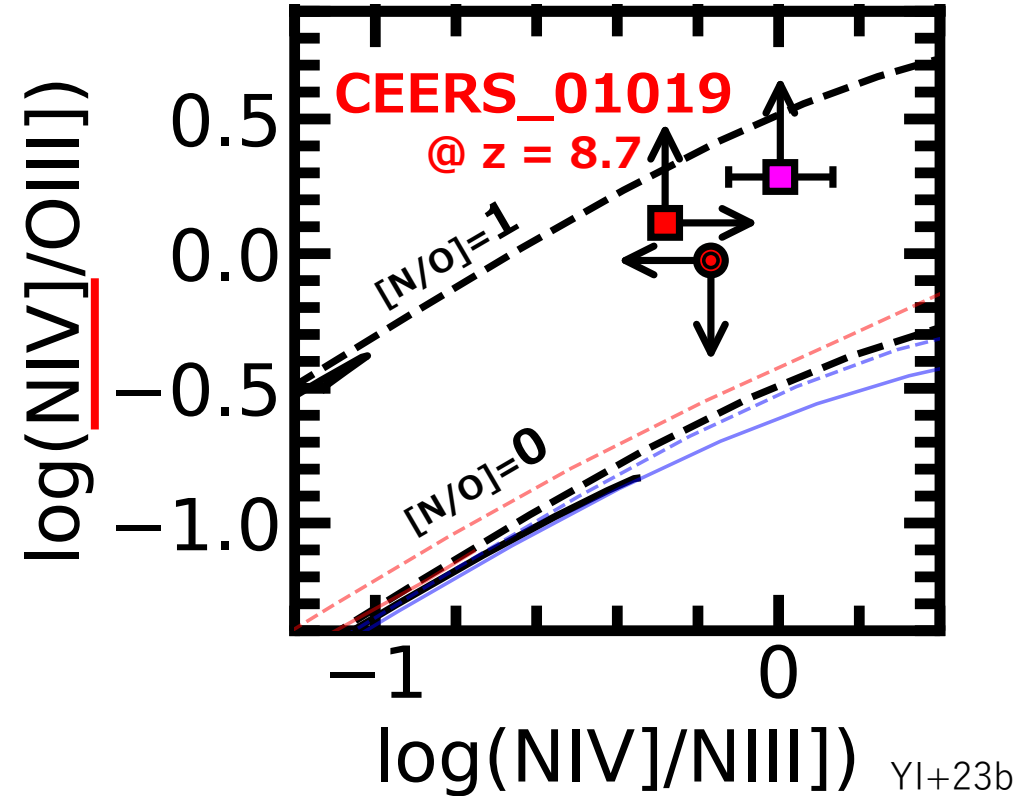
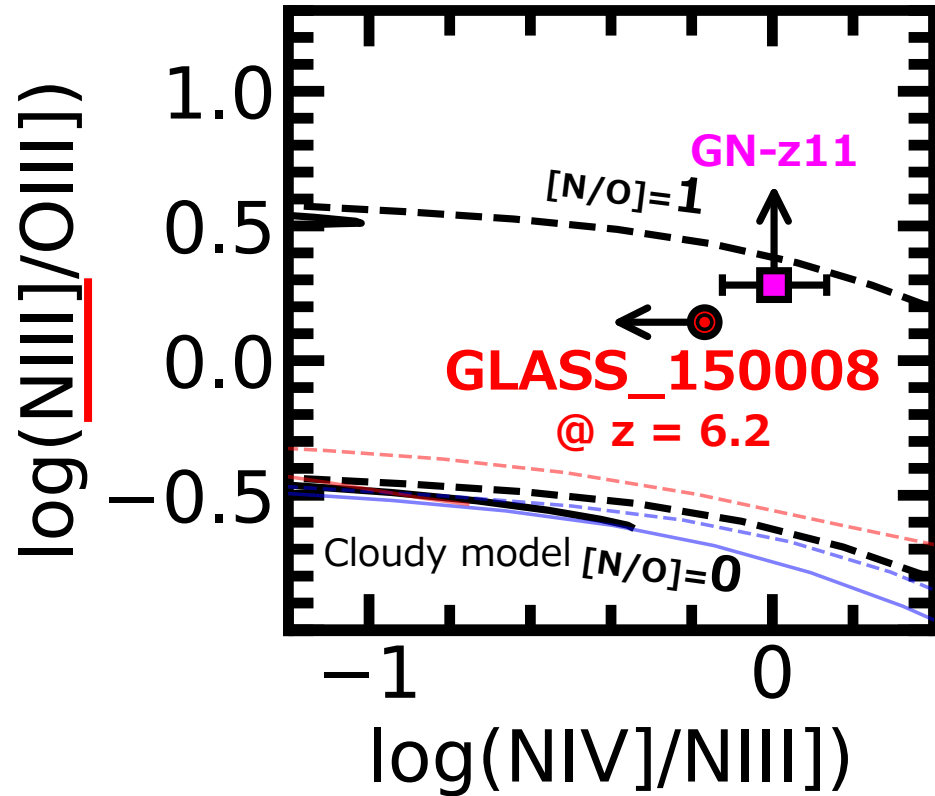


Supersolar N/O (Cameron+23; Senchyna+23)  
→ Any other high- $z$  galaxy w/  $[\text{N}/\text{O}] > 0$ ?



# NIII] or NIV] Detection Other Than GN-z11

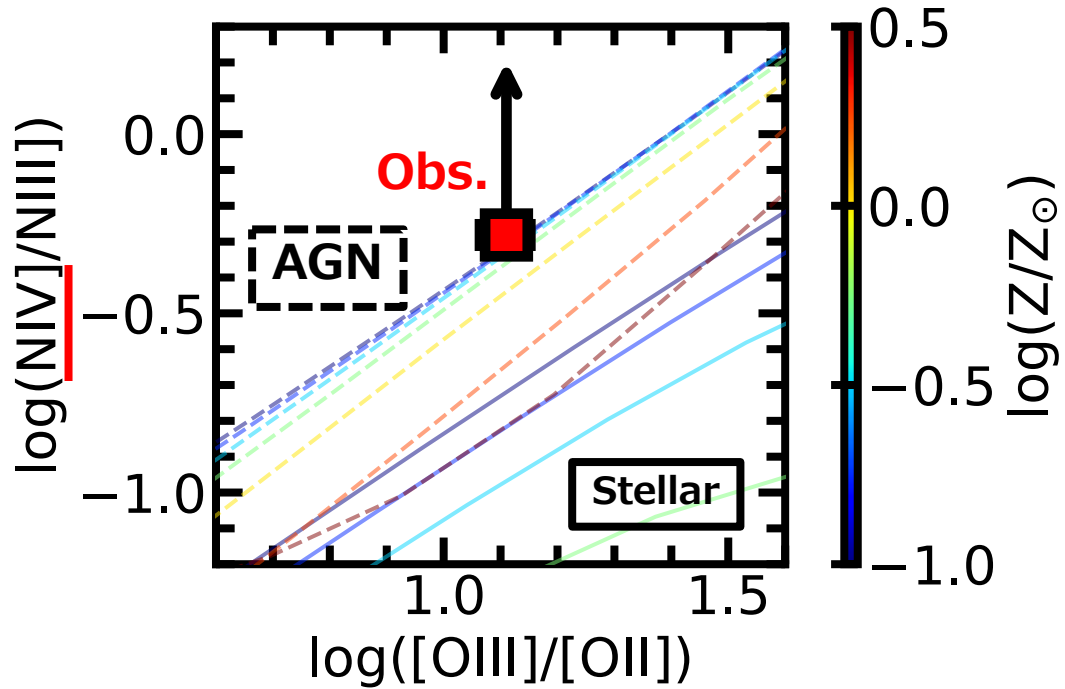
From 70 of Nakajima+23 sample galaxies (w/ grating obs.)



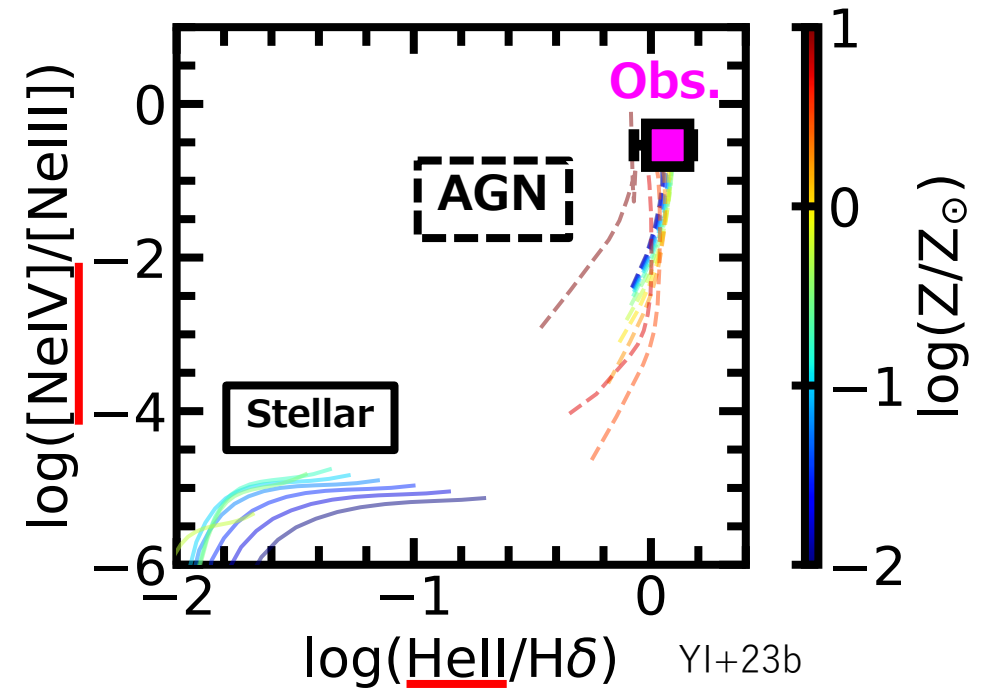
→ Promising to have supersolar N/O

# Star-forming galaxy? AGN?

**CEERS\_01019**

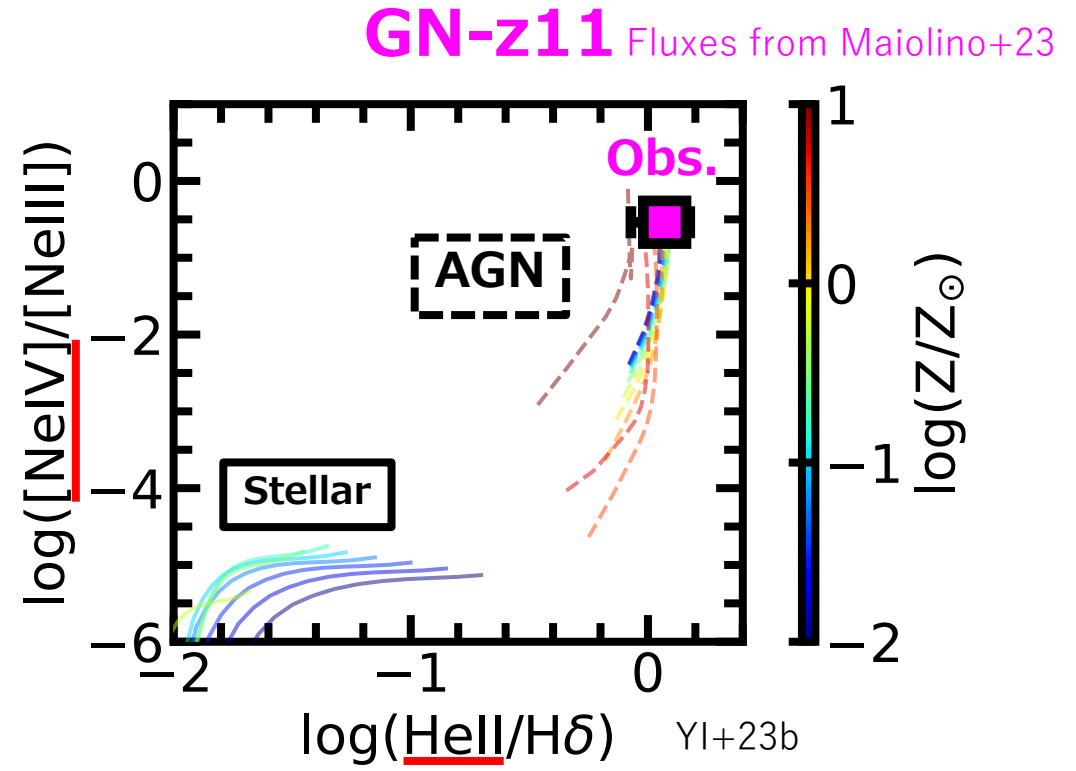
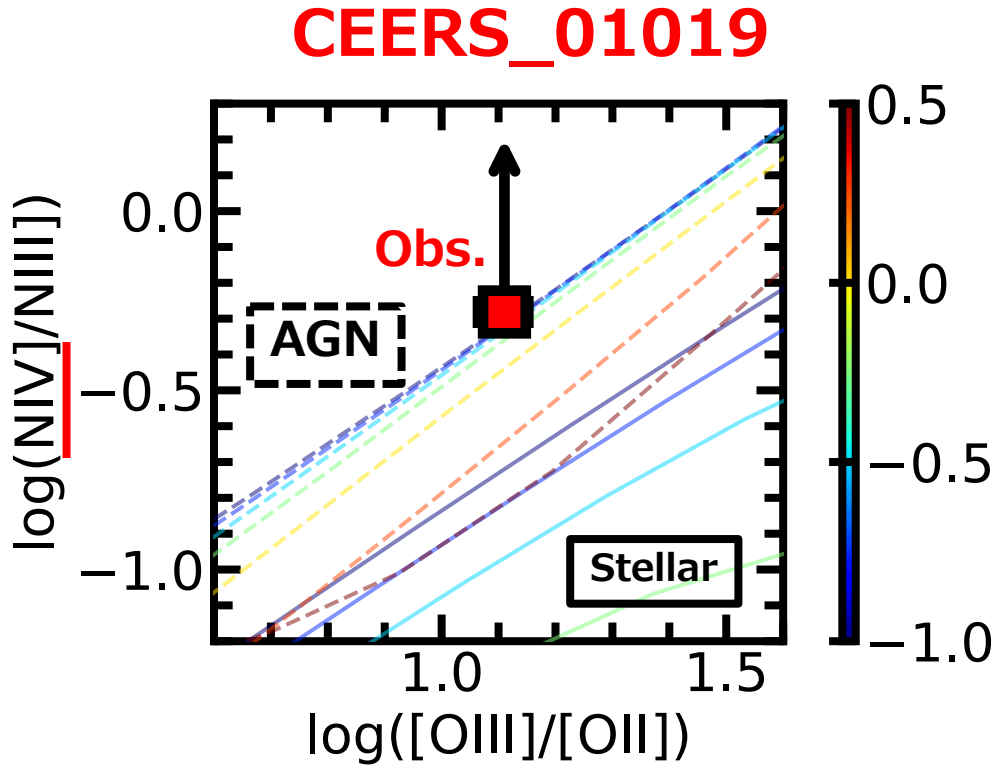


**GN-z11** Fluxes from Maiolino+23



cf. Jan Scholtz's talk

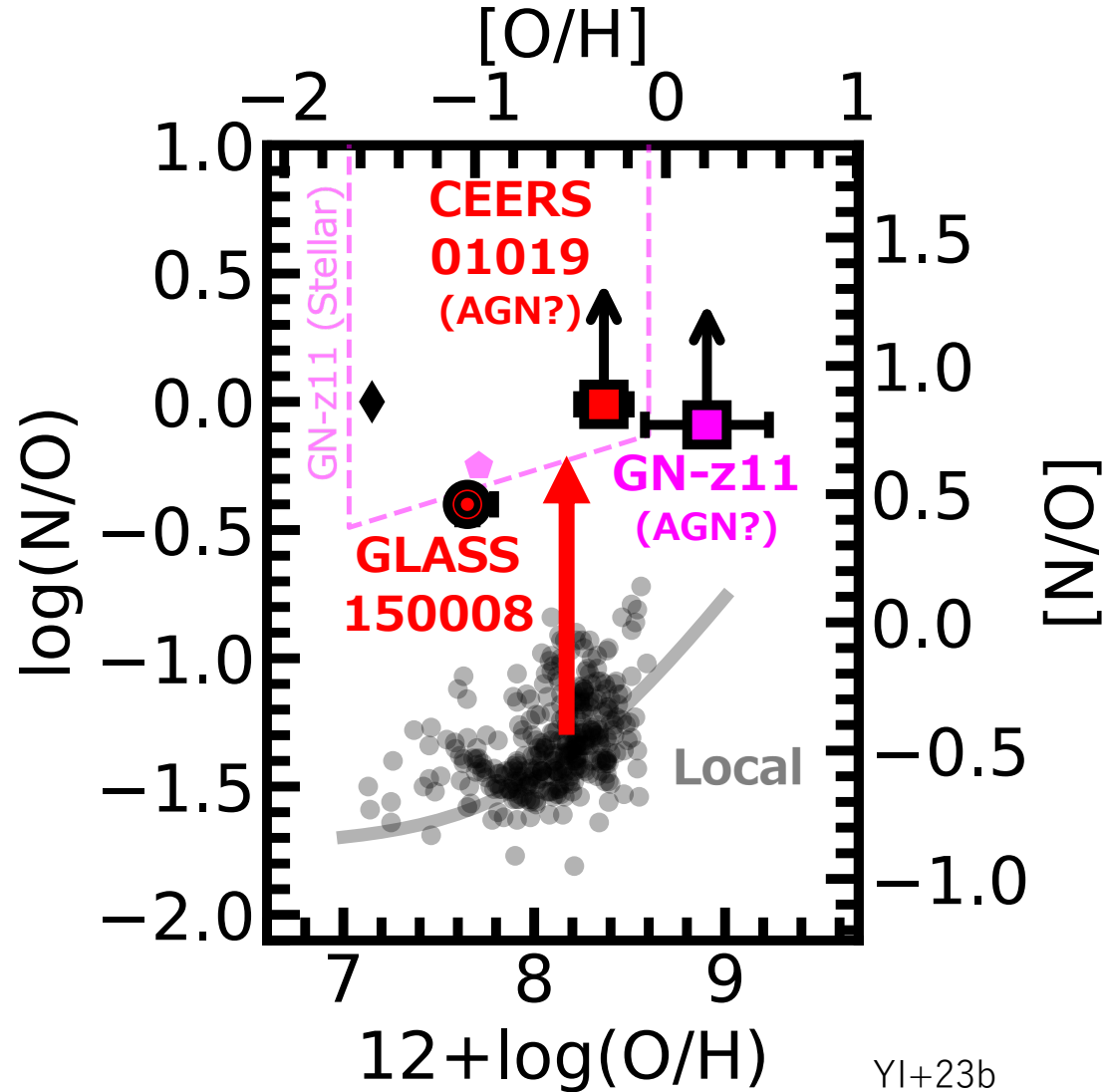
# Star-forming galaxy? AGN?



cf. Jan Scholtz's talk

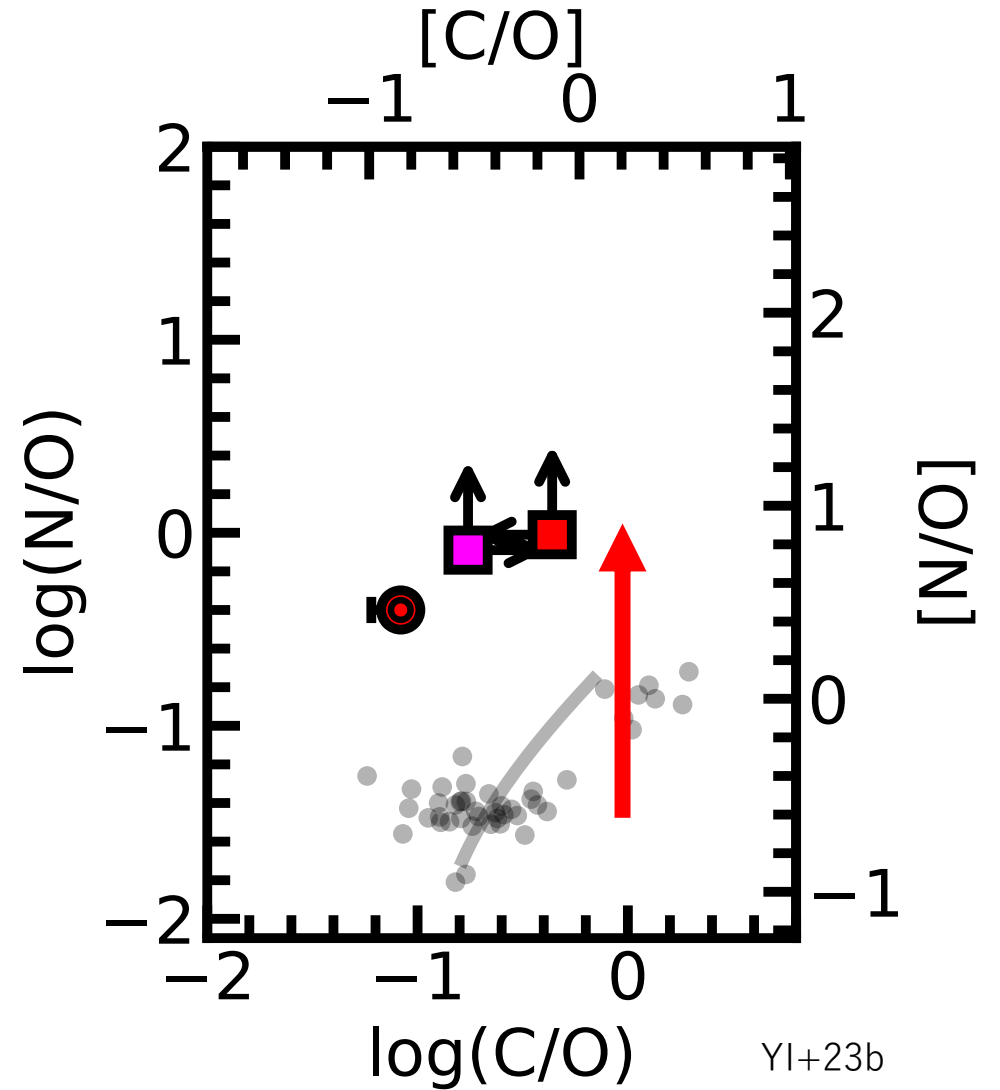
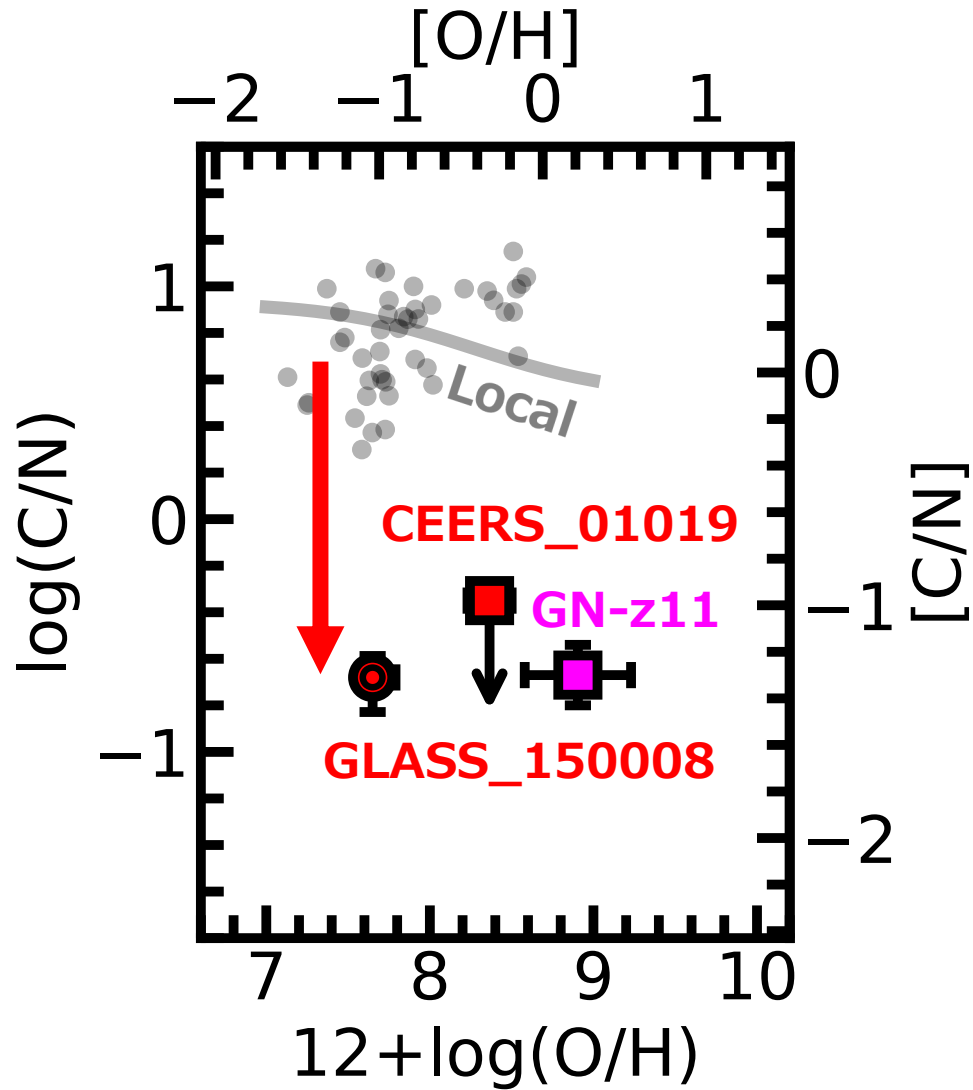
GLASS\_150008 does NOT have high-ionization line  
→ Star-forming galaxy?

# N/O vs. $12+\log(\text{O}/\text{H})$



**$[\text{N}/\text{O}] \gtrsim 0.5$**   
**→ Supersolar N/O galaxies exist other than GN-z11!**

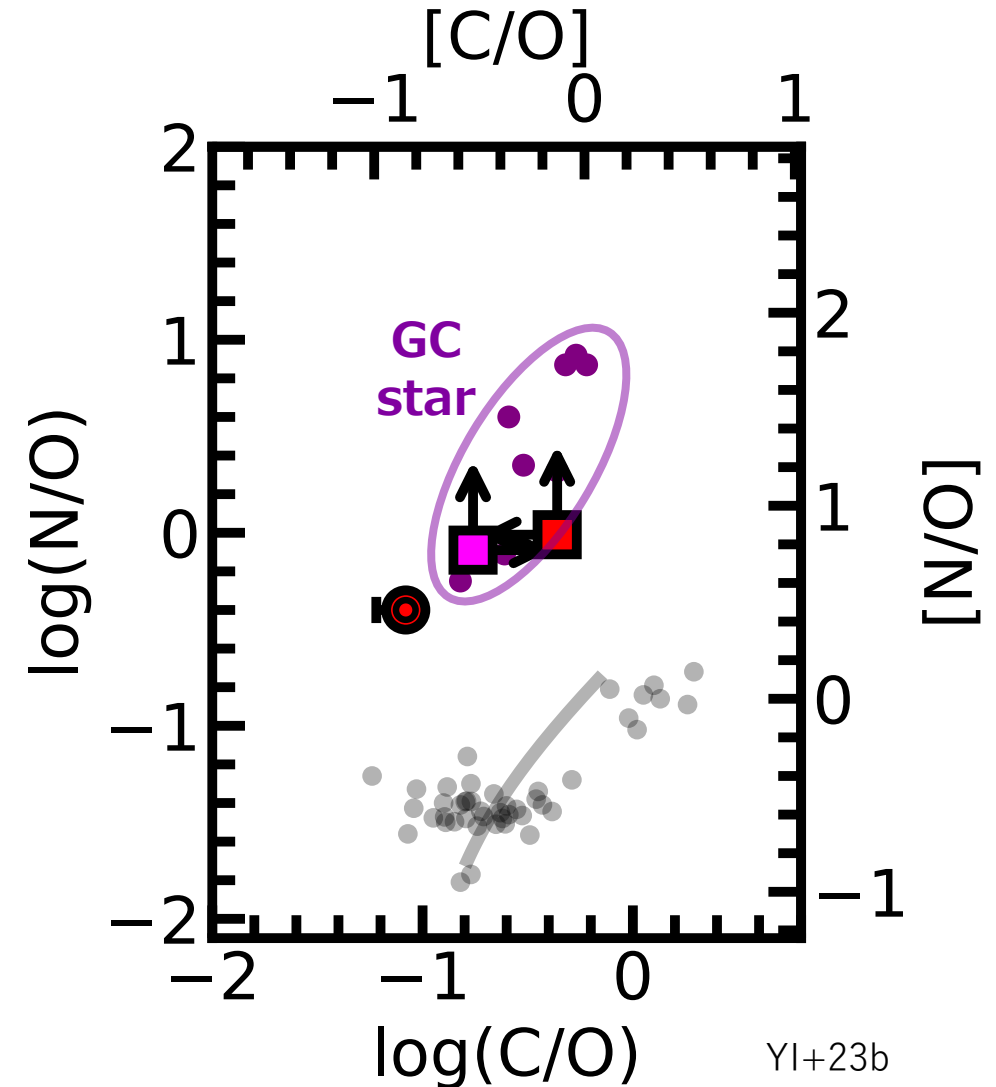
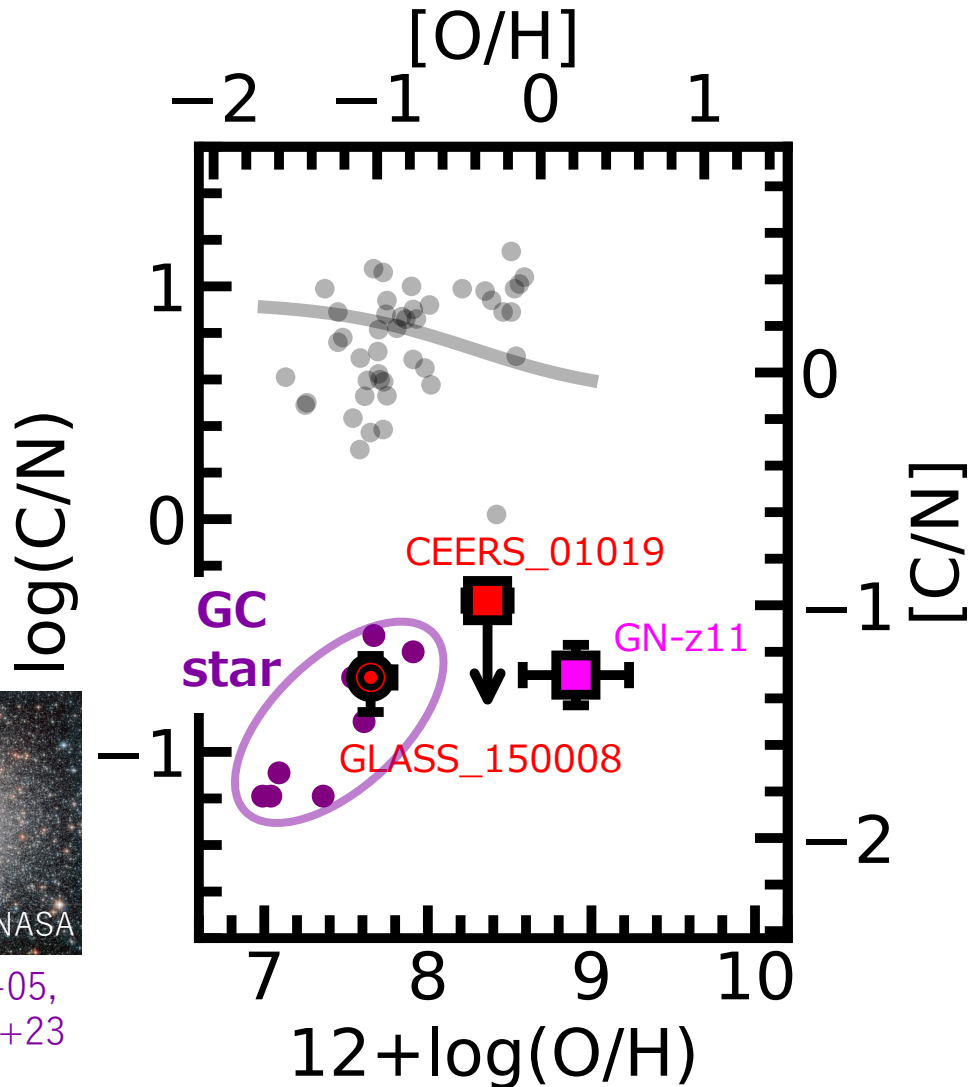
# C/N vs. $12+\log(\text{O}/\text{H})$ & N/O vs. C/O



YI+23b

**→ N selectively enriched**

# Similarity to Globular Cluster (GC) Star

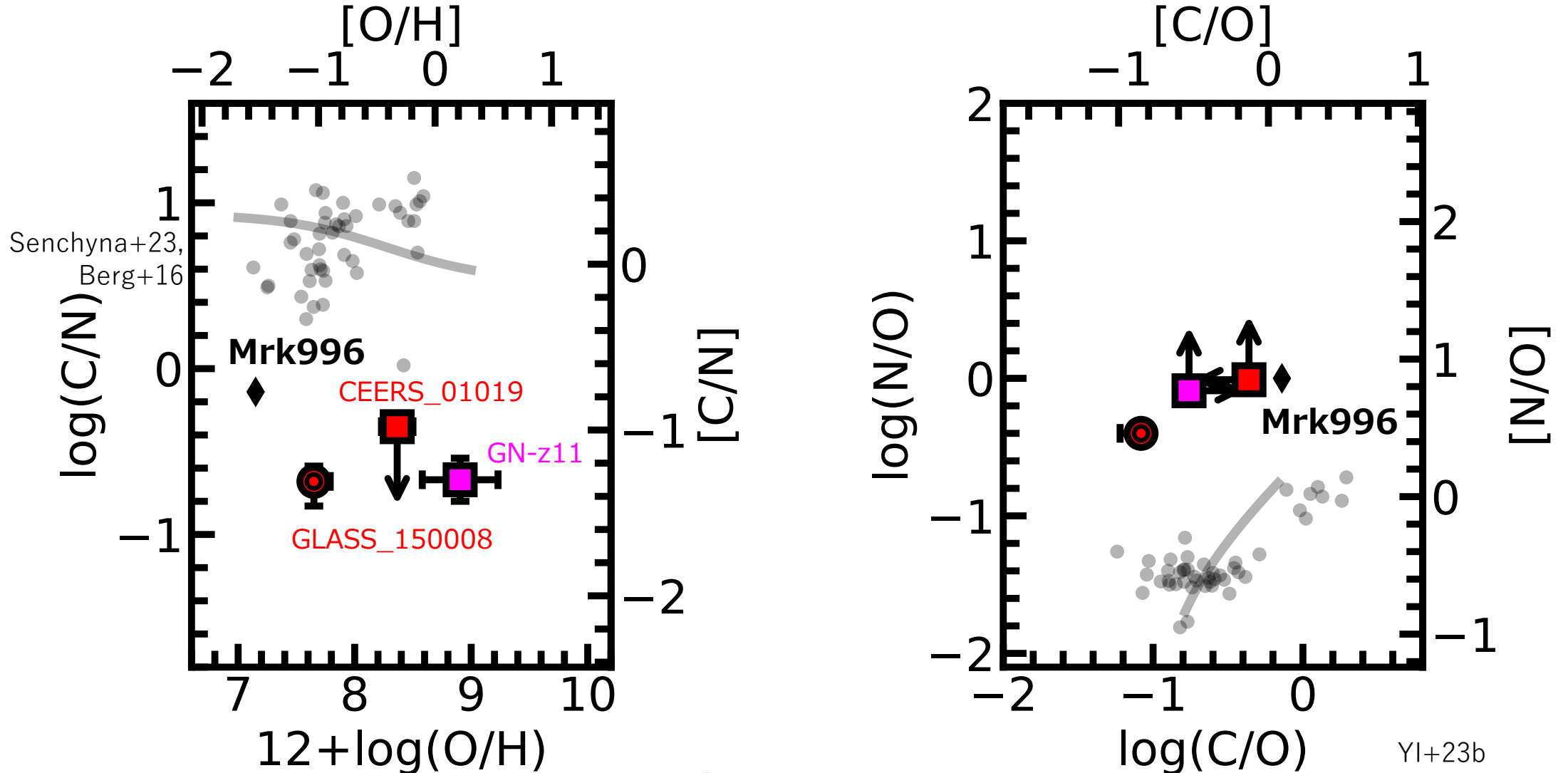


Carretta+05,  
Senchyna+23

→ GC progenitor?

# Similarity to Wolf-Rayet (WR) Galaxy

Mrk996 at  $z \sim 0$ : Hosting WR stars (Telles+14)

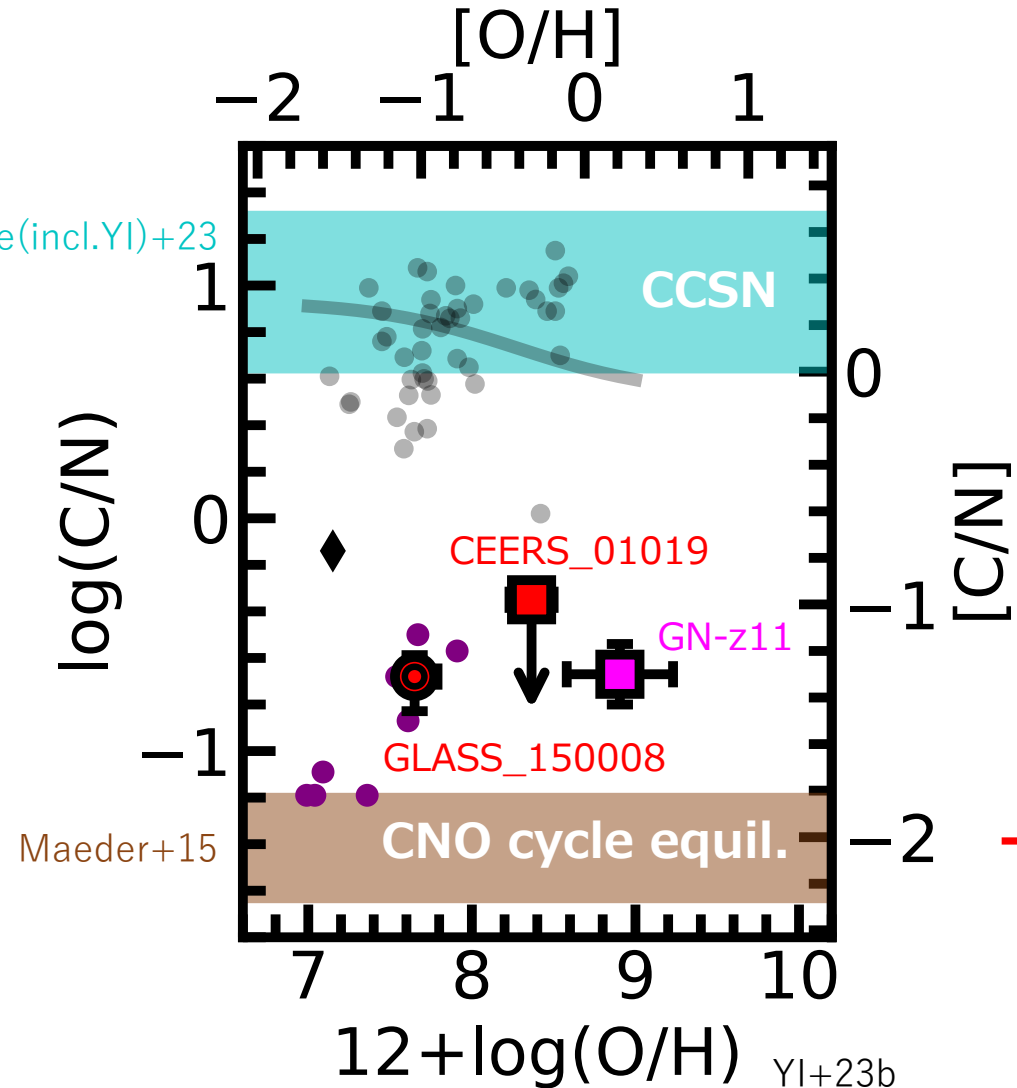


→ Hosting WR stars?

YI+23b

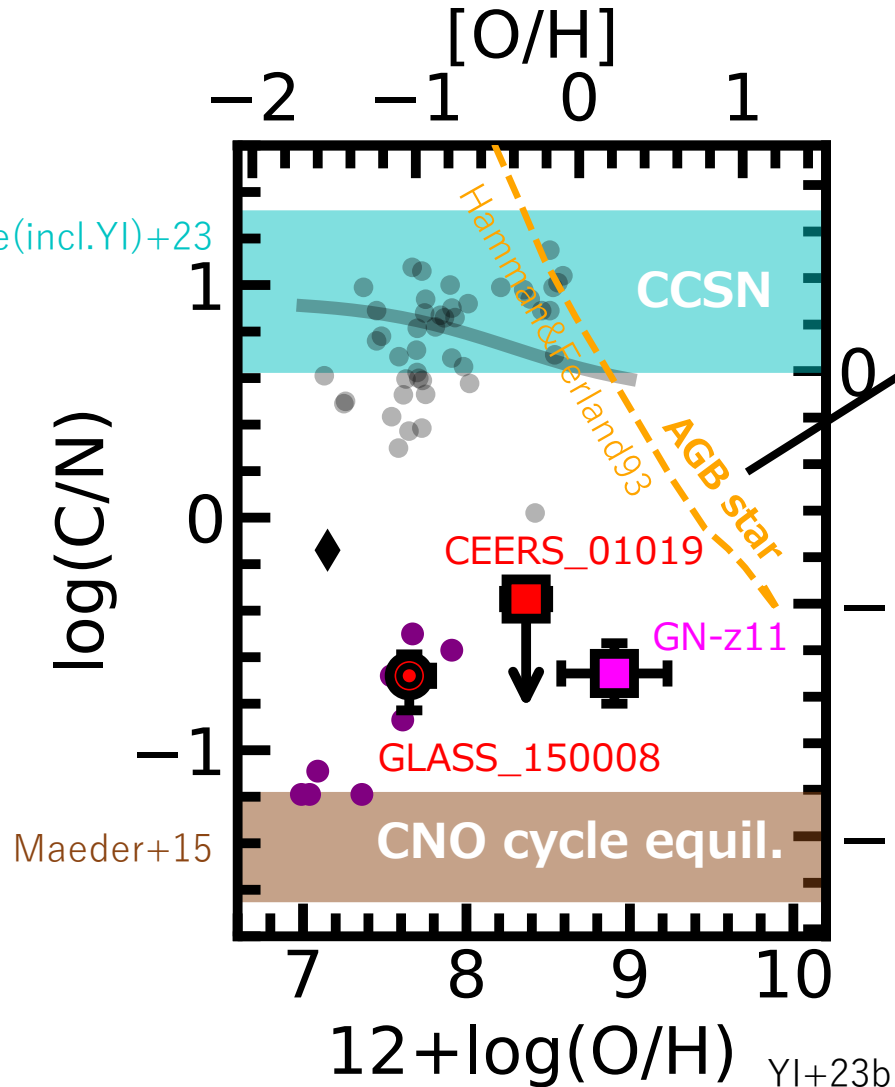


# Origin of Low C/N



→ Dominated by  
CNO-cycle processed gas

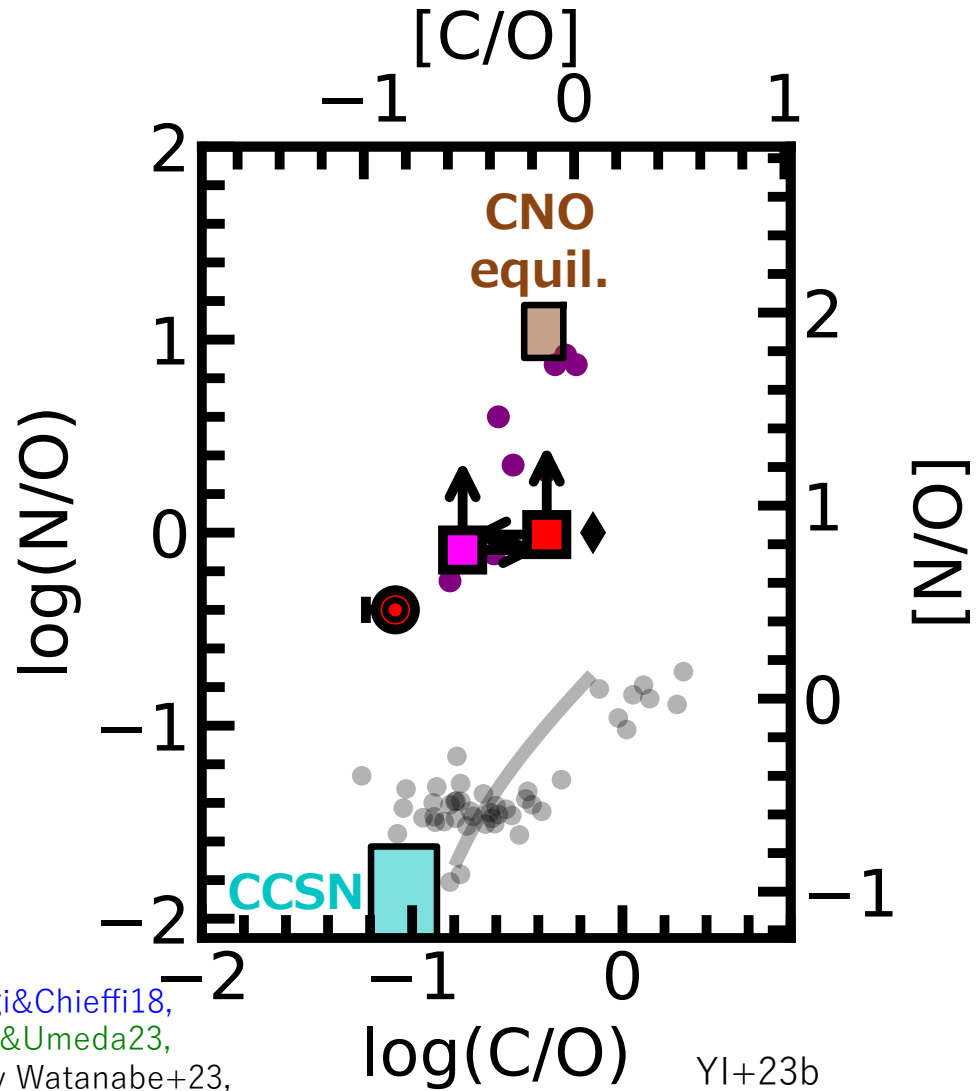
# Origin of Low C/N



(Nitrogen-loud) quasar model incl. **AGB** star  
Too much O/H at  $[C/N] \sim -1$   
**→ Not likely?**

**→ Dominated by  
CNO-cycle processed gas**

# CNO-cycle-based Scenarios

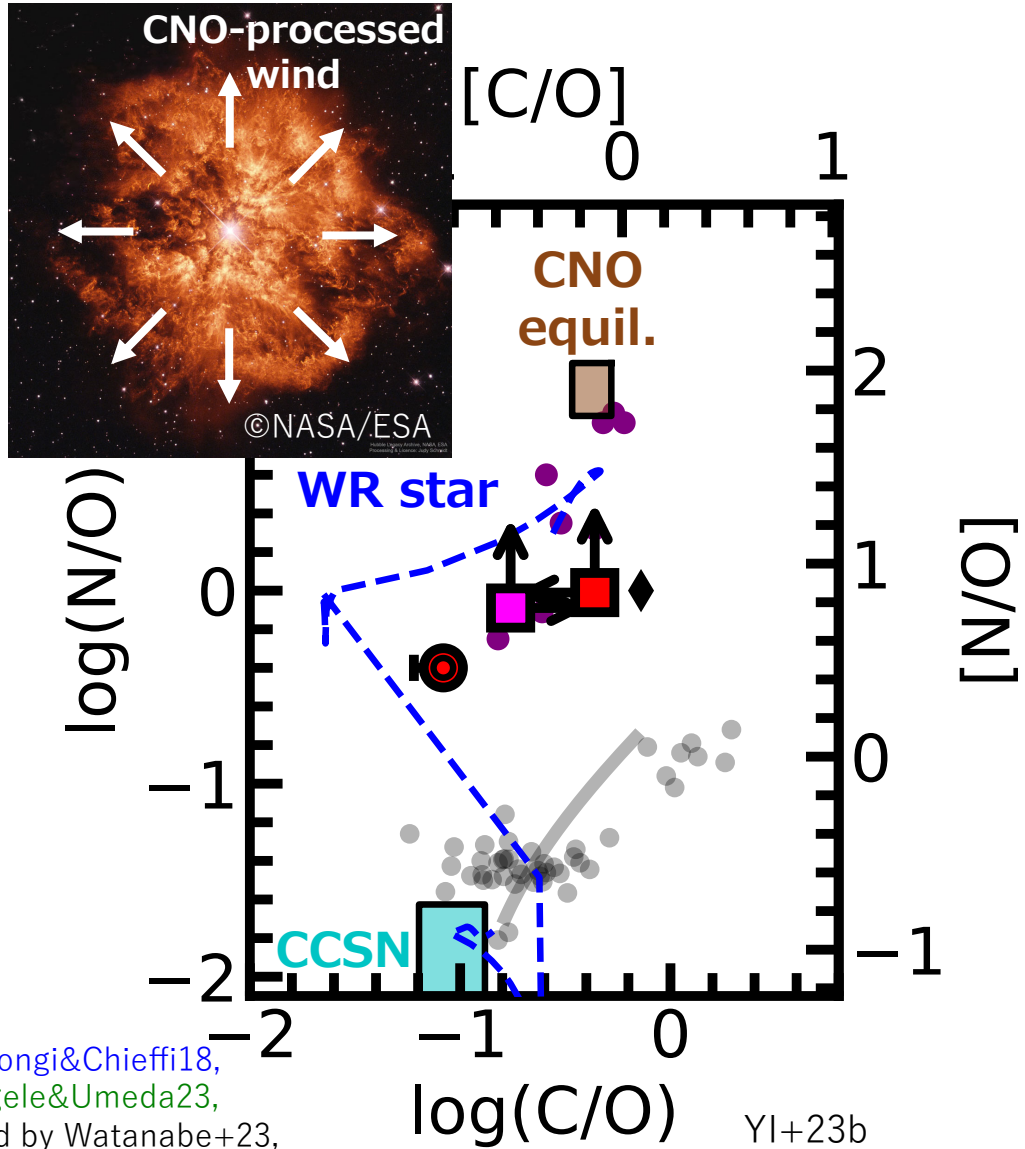


Limongi&Chieffi18,  
Nagele&Umeda23,

Modeled by Watanabe+23,  
Watanabe+in prep

YI+23b

# CNO-cycle-based Scenarios

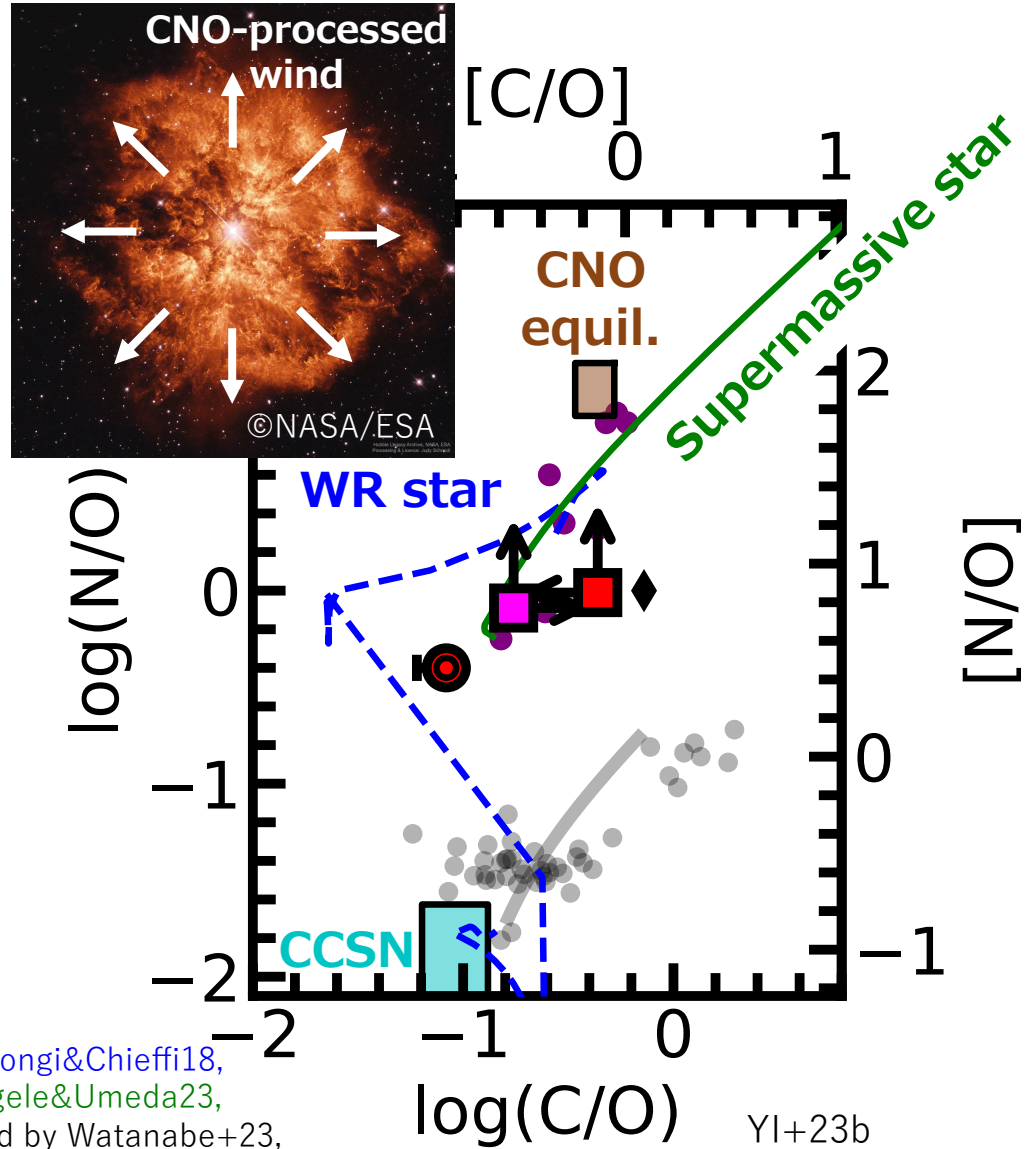


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# CNO-cycle-based Scenarios



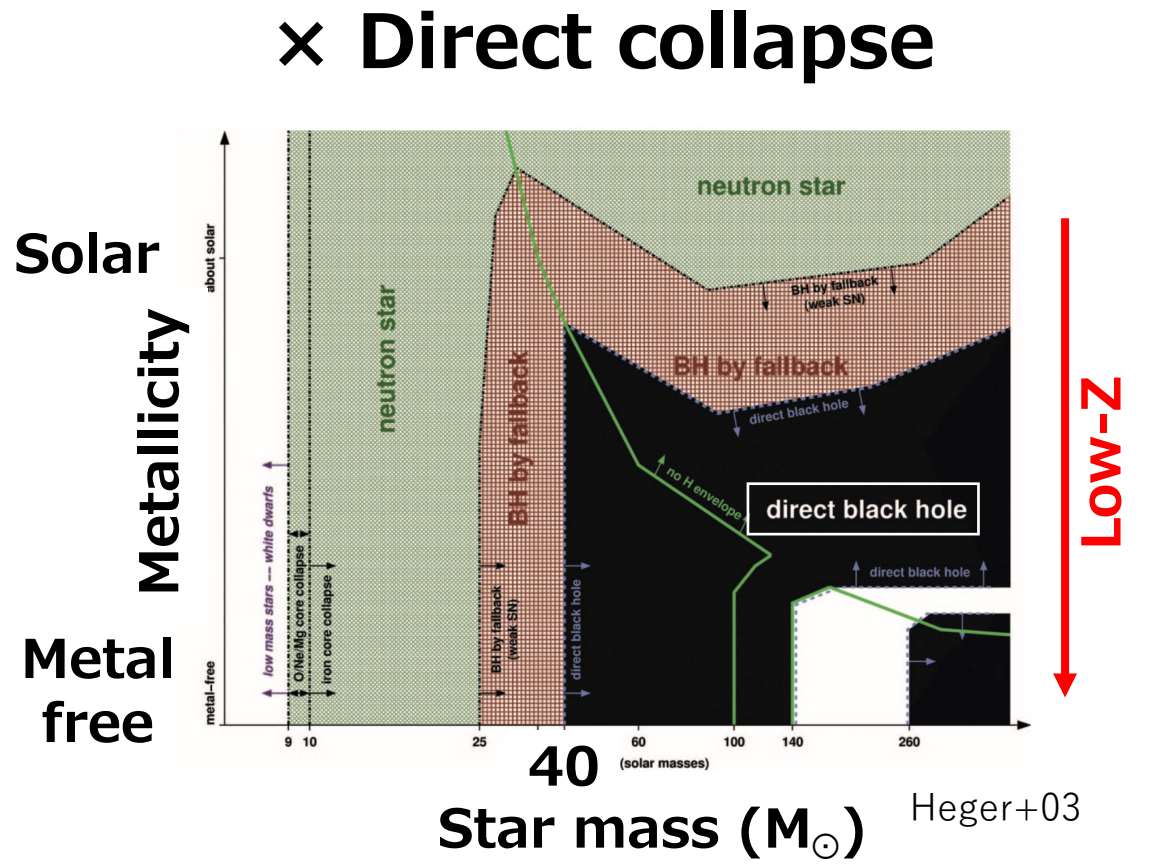
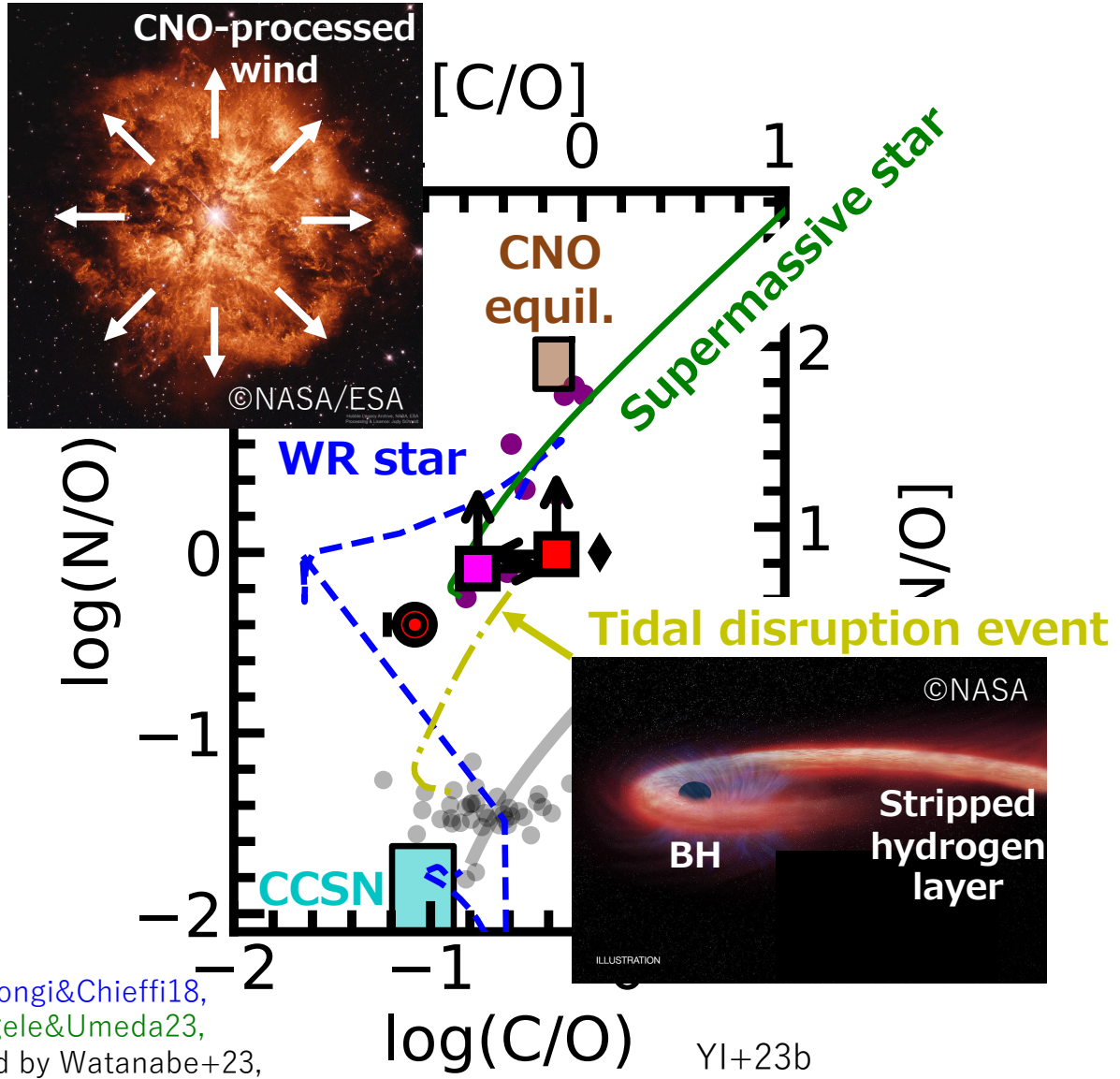
Limongi&Chieffi18,  
Nagele&Umeda23,

Modeled by Watanabe+23,  
Watanabe+in prep

YI+23b



# CNO-cycle-based Scenarios



→ **Low-Z** seems crucial but something more?

Limongi&Chieffi18,  
Nagele&Umeda23,  
Modeled by Watanabe+23,  
Watanabe+in prep

YI+23b

# Summary

**Electron density increases** from  $z \sim 0-1$  to  $1-3$  and  $4-9$

Explainable by a combination of compact morphology and high temperature at high redshifts

Identifying **2** galaxies at  $z > 6$

CEERS\_01019 and GLASS\_150008

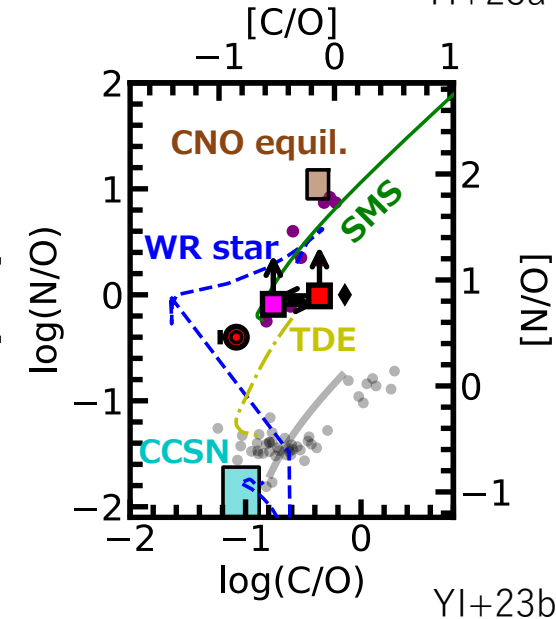
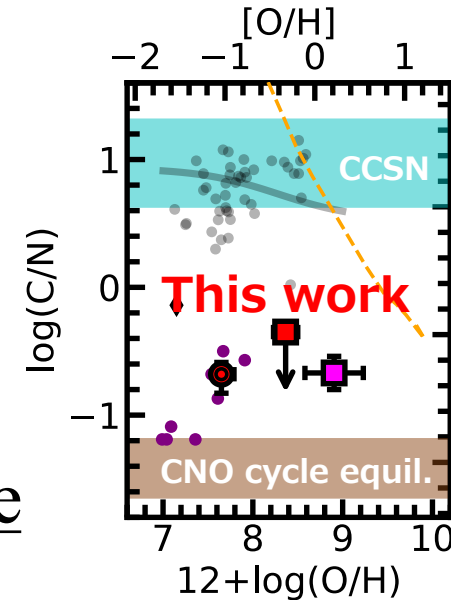
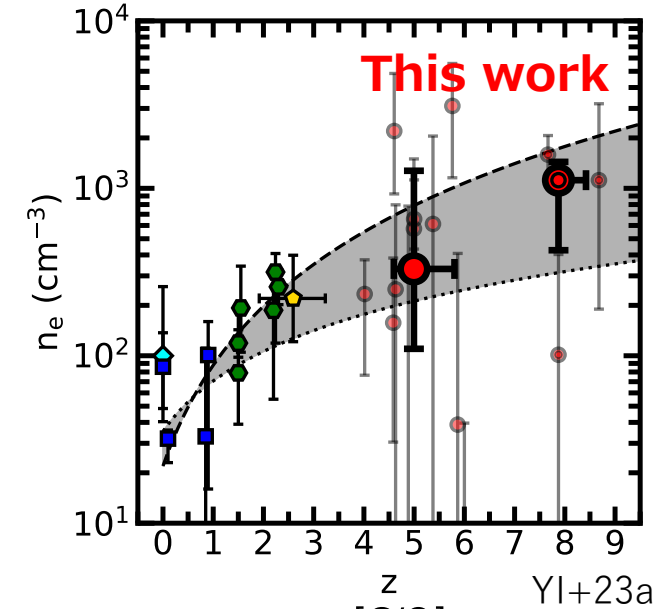
with **supersolar N/O** ratios similar to GN-z11

The 3 galaxies have  **$[C/N] < -1$**

→ Dominated by **CNO-cycle** processed gas

Favoring CNO-cycle-based scenarios

e.g., WR star, SMS, TDE w/ direct collapse



YI+23a

YI+23b